Version



r-Process code for the Java platform



# USER MANUAL

QUARK NOVA PROJECT

# r-Java 2.0 User Manual

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# **Table of Contents**

## Introduction

he rapid capture of neutrons and subsequent beta-decay of the unstable neutron-rich isotopes or rprocess is needed to explain the origin of many heavy neutron-rich nuclei. The astrophysical site where the r-process occurs, however, has not been positively identified. Candidate avenues of research involve the study of winds around proto-neutron stars, ejecta from neutron star mergers and decompressing neutron-rich matter from the surface of neutron stars. Several mechanisms may play a role in this decompression, include the quark nova.

Code has been developed at the Quark Nova Project (QNP) to calculate the abundance of nuclei produced via the r-process. This code is capable of investigating any of the proposed astrophysical sites and as well allows for the study of other possible r-process environments. r-Java is a graphical user interface to this code that makes exploring the parameter space simple and intuitive.

This manual describes the functionality of this graphical interface and assists in the use of the software. The underlying code and methods for the r-process calculation are not described here. For an in-depth review of the code and its application please see the following papers:

"r-Java 2.0 – the nuclear physics". *Mathew Kostka, Nico Koning, Zach Shand, Rachid Ouyed and Prashanth Jaikumar.* 

"r-Java 2.0 – the astrophysics". *Mathew Kostka, Nico Koning, Rachid Ouyed, Prashanth Jaikumar and Zach Shand.* 

These papers are available on the Quark Nova Project (QNP) website at: http://quarknova.ucalgary.ca

## **Quick Start**

-Java 2.0 was designed to be user friendly and easy to use. As such, starting and running a project involves very few steps. This chapter will present you with a quick no-details list of commands to get started. Once started you can explore the interface on your own or read further to understand the full details of the software.

- 1. Launch r-Java 2.0 via java webstart on the QNP website.
- 2. Click on the "Graphs" icon near the center of the main window.
- 3. Click on the graph that appears in order to populate the "Graph Parameters" panel directly to the right of the graph.
- 4. Click on the "X axis" tab in the "Graph Parameters" panel; change the type to "A", set "Min:" to "0", "Max:" to "200" (press "Enter" after you input the values for each) and toggle "Lock" to enabled.
- 5. Click on the "Y axis" tab in the "Graph Parameters" panel; change the type to "Abundance", set "Min:" to "1e-15", "Max:" to "1" (again remember to press "Enter" after you input the values for each), toggle "Lock" and "Log:" to enabled.
- 6. Click on the "Code" tab in the right-most panel of the main window and choose "NSE" (Nuclear Statistical Equilibrium) as the code type.
- 7. Enter the following values (remember to press "Enter" after you input the values for each): T = 3 E9,  $\rho = 1E7$ , Ye = 0.3 and click on the "Calculator" icon in the main toolbar. The abundances will then be calculated and displayed on the graph.
- 8. Select "r-Process" from the "Type:" drop-down box, located at the top of the right-most panel of the main window.
- 9. Switch to the "Data" module by clicking on the icon along the main toolbar. Click the "Restore" button (green curved arrow) at the top of the data table. **Note:** It will take a few seconds to restore the data to default values, so if you move on to the next step and the scroll bar does not respond, that is why...just wait a couple more seconds...there it should be done now.

Running NSE

- 10. Scroll down in the table to "Element" = Fe, "Z" = 26, "A" = 56. In this row, under the column "Initial MF" change the value from 0.0 to 0.3 (remember to press "Enter" after making this change).
- 11. In the "Code Parameter" panel, enter the following parameter values (remember to press "Enter" after each entry):  $T_0 = 1E9$ ,  $\rho_0 = 1E8$ ,  $\tau = 0.001$ , Decay time = 1E4 and Update = 50 and click on the "Calculator" icon in the main toolbar. The abundances will then be calculated and displayed on the graph. You can navigate between the "Messages" module (to view the evolution of the physical parameters) and the "Graphing" module (to view the changes in the nuclei abundances) without disturbing the code.
- 12. Select "Fission" from the "Type:" drop-down box, located at the top of the right-most panel of the main window.
- 13. Click on the blue plus symbol at the bottom of the right-most panel of the main window. Select "Z: 92.0 A: 235.0" from the list shown in the pop-up window and click "OK".
- 14. Click on the "Calculator" icon in the main toolbar. The fission fragmentation distribution will be calculated and displayed on the graph.
- 15. To store your results for comparison with a future simulation click on the "Data" tab in the "Graph Parameters" panel (directly to the right of the graph). Click on the words "Primary Data [Point]" in the text field and then in the "Properties" panel that appears below change the "Name:" and "Color:" fields. The next simulation run will regenerate a "Primary Data [Point]" data set.
- 16. To save the graph, right click anywhere on the graph and choose "Save Ascii" or "Save Image" depending on if you want to save the text data, or the image respectively.
- 17. Save the project by clicking the "Save" icon in the main toolbar or by clicking on the "Desktop" icon on the main toolbar (left-most icon) and then clicking on the "Save" or "Save As" icons.
- 18. You can then open this project at a future time by using the "Open" icon in the toolbar, or the "Open" icon on the "Desktop" or if the project appears under "Recent projects" on the "Desktop" by clicking on the icon with the desired project name.

The above steps should get you started in exploring the user interface. Feel free to change parameters and options to see what the program can do.

Running fission

# 3

## **Graphical User Interface**

-Java is graphical user interface to the r-process code. Navigating the toolbars, menus and graphs is meant to be intuitive. This section will briefly cover the function of each interface component. An important point to note is that when you enter a value in a text box, you must press "Enter" for the change to be recorded. r-Java 2.0 does not monitor key presses and will not update unless "Enter" is pressed. Individual modules can be detached from the main r-Java 2.0 window by clicking the 💌 icon.

## Toolbar



Fig. 1: The toolbar.

The toolbar is always present along the top of the main window of r-Java 2.0. The toolbar provides a graphical interface to some common commands in r-Java 2.0. The icons represent:

- **Desktop Icon**: This icon sends the user to the desktop module.
- Calculate Icon: This is the most important icon of r-Java 2.0 which starts off any abundance calculation.
- **Open Icon**: Opens a saved project.
- **Save Icon**: Saves the current project.
- **Graphs Icon**: This icon sends the user to the graphing module.
- Elements Icon: This icon sends the user to the 3D periodic table module.
- Messages Icon: This icon sends the user to the messages module.

• Data Icon: This icon sends the user to the data module.

## Desktop

The desktop is the first screen that users will be greeted with when running r-Java 2.0. The desktop contains quick links to the different modules of r-Java 2.0 as well as system tools.

#### **System Tools**



**New:** This creates a new, blank project. When r-java 2.0 is first started, a new project is automatically created. This menu item is useful if you have a project loaded and you want to start a new one.





**Open**: Opens a saved project.

**Desktop:** Quick link to the desktop.

Fig. 2: The desktop.

**Save**: Saves the current project to the current filename of the project. If the project has not yet been saved, a dialog will be displayed asking for the name of the file you wish to save to.



Save As: Saves the project to a filename of your choice.

#### **Modules**



**Graphs:** Quick link to the graphing module. As well as displaying the graphs this is the module where the user can output the calculated abundances.



**Messages**: Quick link to the messages module. This module displays updates on the physical parameters of the system (such as temperature and density) at the current simulation time. Error messages are as well displayed in this module.



**Data**: Quick link to the data module. This module displays all the nuclear data (such as masses and reaction rates) for each nucleus in the network. With this module the user is able to adjust the nuclear data for any or all of the nuclei.



Elements: Quick link to the 3D periodic table module.

#### Commands



#### Information



**Memory**: Opens a window that displays memory information such as; average memory usage, total memory used and memory usage rate.



Progress: Opens window that displays the progress of the current simulation run.



Help: Opens the user manual.

#### **Recent projects**

In the "recent projects" section of the desktop a list of up to six (6) of the most recently used projects are displayed for quick access.



Project name: Loads the r-Java 2.0 project of the given name.

## **Data Handling**

Since r –Java 2.0 calculates abundances of nuclei, certain nuclear data, such as neutron capture cross-sections and decay rates, is needed before the calculations can commence. By default r-Java 2.0 has a set of all the necessary data loaded, see the paper "*r-Java 2.0 – the nuclear physics*" for details on the derivation of the data.

All of the data used by r-Java 2.0 can be viewed in the "Data module". This module is dominated by an editable table that contains all of the data.

The user can adjust the amount of data displayed by toggling on/off the selections listed in the "Display" panel. Regardless of these selections the table will always include; "**Element**", "**Z**" (number of protons) and "**A**" (atomic mass number).

-	-	-							
2	1	<b>C</b>							
Refresh	Clear	Restore							
Element	z	A	N	Mass [amu]	Solar MF	Mass Fracti.	Initial MF	Alpha Dec	Parameters
Co	27.0	81.0	54.0	81.0029	0.0	1.5827860	0.0	0.0	Display
Co	27.0	82.0	55.0	82.0113	0.0	7.4513932	0.0	0.0	General Data:
Co	27.0	83.0	56.0	83.0179	0.0	7.5206941	0.0	0.0	Nuclear Balance III
Co	27.0	84.0	57.0	84.0266	0.0	5.4180428	0.0	0.0	Nuclear Data:
Co	27.0	85.0	58.0	85.0339	0.0	2.2693440	0.0	0.0	Fission Data:
Co	27.0	86.0	59.0	86.0427	0.0	1.6357599	0.0	0.0	(n,y)(y,n) Data:
Co	27.0	07.0	60.0	07.0504	0.0	3.4970074	0.0	0.0	
Co	27.0	88.0	61.0	88.0599	0.0	3.7883740	0.0	0.0	Batch Data Adjustments
Co	27.0	89.0	62.0	89.0678	0.0	9.0316740	0.0	0.0	Preset Data
Co	27.0	90.0	63.0	90.0776	0.0	0.0	0.0	0.0	Mass Model: HFB21
Co	27.0	91.0	64.0	91.0859	0.0	0.0	0.0	0.0	
Co	27.0	92.0	65.0	92.0953	0.0	0.0	0.0	0.0	Custom Data
Ni	28.0	46.0	18.0	46.0491	0.0	0.0	0.0	0.0	O Data type: Nuclear V
Ni	28.0	47.0	19.0	47.0358	0.0	0.0	0.0	0.0	
Ni	28.0	48.0	20.0	48.0195	0.0	0.0	0.0	0.0	
Ni	28.0	49.0	21.0	49.0096	0.0	0.0	0.0	0.0	Column Separator
Ni	28.0	50.0	22.0	49.9953	0.0	0.0	0.0	0.0	C Tab
Ni	28.0	51.0	23.0	50.9885	0.0	0.0	0.0	0.0	Space
Ni	28.0	52.0	24.0	51.9755	0.0	0.0	0.0	0.0	Other
Ni	28.0	53.0	25.0	52.9692	0.0	0.0	0.0	0.0	
Ni	28.0	54.0	26.0	53.9579	0.0	0.0	0.0	0.0	Export
Ni	28.0	55.0	27.0	54.9523	0.0	0.0	0.0	0.0	
Ni	28.0	56.0	28.0	55.9433	0.0	0.0	0.0	0.0	Import
Ni	28.0	57.0	29.0	56.9403	0.0	0.0	0.0	0.0	L'
Ni	28.0	58.0	30.0	57.9355	4.94E-5	0.0	0.0	0.0	Import Initial Mass Fraction
Ni	28.0	59.0	31.0	58.935	0.0	0.0	0.0	0.0	Filename.
Ni	28.0	60.0	32.0	59.9314	1.96E-5	0.0	0.0	0.0	
Ni	28.0	61.0	33.0	60.9317	8.59E-7	0.0	0.0	0.0	Column Separator
Ni	28.0	62.0	34.0	61.9291	2.78E-0	0.0	0.0	0.0	O Tab
Ni	28.0	63.0	35.0	62.9306	0.0	0.0	0.0	0.0	
Ni	28.0	64.0	36.0	63.9283	7.27E-7	0.0	0.0	0.0	Space
Ni	28.0	65.0	37.0	64.9306	0.0	0.0	0.0	0.0	Other
Ni	28.0	66.0	38.0	65.9292	0.0	0.0	0.0	0.0	- Francisco -
Ni	28.0	67.0	39.0	66.9321	0.0	0.0	0.0	0.0	Import
Ni	28.0	68.0	40.0	67.9322	0.0	0.0	0.0	0.0	
Ni	28.0	69.0	410	68 9357	0.0	0.0	0.0	0.0	

Fig. 3: The data module.

- <u>General Data</u>: When selected the table will include; "Mass [amu]" (the atomic mass of the nuclei), "Solar MF" (the solar mass fraction of the nuclei), "Mass Fraction" (the computed mass fraction of the nuclei for the last simulation run) and "Initial MF" (used in the full reaction network calculation as the initial abundance of the r-process simulation).
- <u>Nuclear Data:</u> When selected the table will include; "Alpha Decay Rate [s^-1]", "Beta Decay Rate [s^-1]", "Beta Decay Q value" (the energy released by a beta decay), "Prob. 0 bdn's" (the probability that during beta decay one neutron will be emitted), "Prob. 2 bdn's" (the probability that during beta decay two neutrons will be emitted), "Prob. 3 bdn's" (the probability that during beta decay three neutrons will be emitted). The

General Data:	~	
Nuclear Data:	~	
Fission Data:	~	
(n,y)(y,n) Data:	~	

probability of beta-delayed neutron emission can be anything between 0 and 1 however r-Java 2.0 will automatically normalize the probabilities after any change.

- Fission Data: When selected the table will include; "Prob (BDF)" (the probability that beta decay will lead to immediate fission of the resultant nucleus), "Fission Barrier [MeV]" and a list of temperature dependent neutron-induced fission rates, by default there a 28 different temperature points denoted as "nf\_T0",...,"nf\_T27". The temperature grid points are (in units of 10<sup>9</sup> K); 0.001, 0.005, 0.01, 0.05, 0.1, 0.15, 0.2, 0.25, 0.3, 0.4, 0.5, 0.6, 0.7, 0.8, 0.9, 1, 1.5, 2, 2.5, 3, 3.5, 4, 5, 6, 7, 8, 9, 10. r-Java 2.0 uses a custom interpolation technique in order to determine the intermediate rates.
- $(n, \gamma)(\gamma, n)$  Data: When selected the table will include a list of the Maxwellian-averaged neutron capture cross sections ( $N_A < \sigma \nu >$  in units of mol<sup>-1</sup> cm<sup>3</sup> s<sup>-1</sup>) and ( $\gamma$ , n) Maxwellian-averaged photo-rates (in units of s<sup>-1</sup>). The temperature grid points are the same as that for the neutron-induced fission rates (in units of 10<sup>9</sup> K); 0.001, 0.005, 0.01, 0.05, 0.1, 0.15, 0.2, 0.25, 0.3, 0.4, 0.5, 0.6, 0.7, 0.8, 0.9, 1, 1.5, 2, 2.5, 3, 3.5, 4, 5, 6, 7, 8, 9, 10.

## **Adjusting Data**

The user can change any adjustable parameter by double-clicking its cell in the table located in the Data module. The parameters that are <u>not adjustable</u> are; **Z**, **A**, **N**, **Solar Mass Fraction (MF) and Mass Fraction** (this column is where the output of the calculation is displayed). If the user is interested in adjusting a large set of data using the table would be cumbersome and thus r-Java 2.0 provides the user with the option to change the data as a batch.

#### **Batch Data Adjustments**

There is an option in r-Java 2.0 to simply switch between nuclear mass models. By switching the mass model a new set of nuclei is loaded into r-Java 2.0 with masses and associated rates calculated with the logic of the chosen mass model.

If the user of r-Java 2.0 chooses to make custom changes to the data **we recommend that the user first exports the data set** that will be changed in order to preserve the scope of the network. Then **once the user has made changes to the data set, import the data** back into r-Java 2.0. This is recommended because typos such as making an isotopic chain discontinuous could lead to unrealistic results from the simulations. As well first exporting the data and then modifying

Batch Data Adjustmer	nts
Preset Data	
O Mass Model:	HFB21
Custom Data	
Data type:	Nuclear
Column Sepa	rator
🔘 Tab	
🔘 Space	
Other	
Export	
Import	
Fig. 5: Batch data ad-	iustments.

the output file ensures that the format is correct when re-importing the data set.

• Changing Temperature Dependent Rates: The neutron-induced fission cross-sections, the neutron capture cross-sections and photorates are interpolated from a temperature grid. By default the temperature grids have 28 points (in units of 10° K; 0.001, 0.005, 0.01, 0.05, 0.1, 0.15, 0.2, 0.25, 0.3, 0.4, 0.5, 0.6, 0.7, 0.8, 0.9, 1, 1.5, 2, 2.5, 3, 3.5, 4, 5, 6, 7, 8, 9, 10) and each nuclei has a rate (and cross-sections) associated with each grid point. If the user wishes to change the range of the temperature grid this can be done while importing a new data set. An excerpt of a  $(n,\gamma)(\gamma,n)$  data file can be seen in Fig. 6. The series of numbers (separated by tabs) following the "T9" defines the range of the temperature grid (in units of 10° K), the numbers must be increasing from left to right.

🥘 rjIn	put_r	nggn.txt -	Notepad										
<u>F</u> ile [	Edit	F <u>o</u> rmat	<u>V</u> iew <u>H</u> elp										
т9		1.00E-	·03	5.00E-03	1.00E	-02	5.00E-02	2	1.00E-0	1 1.	50E-01	2.00E	-01 🔺
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Ő		1	0.0000	00E+00	0.00000E+00	0.00	0000E+00	0.00000	0E+00	0.00000E+	00 0.0	00000E+00	0.
L		1 2	4.45330	00E+04 20E+02	4.415190E+04 2.974240E+02	4.38	3750E+04	4.125/8	0E+04 0E+02	3.8/3/50E+ 5.125720E+	04 3.6 02 5.9	/6100E+04 85710E+02	3. 6.
		3	0.0000	00E+00	0.00000E+00	0.00	0000E+00	0.00000	0E+00	0.00000E+	0.0	00000E+00	0.1
2		4	0.00000	10E+02 00E+00	0.000000E+00	9.38	00000E+02	0.00000	0E+03	0.000000E+	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	00000E+00	0.1
<b>F</b> '	ζ.Τ	- 1		. (71									

For the case of the  $(n,\gamma)(\gamma,n)$  data file each row after the commented out section (comment string is #) follows the same format; Z, A,  $N_A < \sigma v > @$  T0,  $N_A < \sigma v > @$  T1,...,  $N_A < \sigma v > @$  TN,  $\lambda_{\gamma} @$  TN,  $\lambda_{\gamma} @$  T0,  $\lambda_{\gamma} @$  T1,...,  $\lambda_{\gamma} @$  TN, where T0 represents the first temperature point in the grid and TN the last temperature point in the grid. When making changes to the  $(n,\gamma)(\gamma,n)$  data file the user must ensure that there is a neutron capture cross-section and photorate associated with every temperature grid point.

## Codes

sotope abundances are the main output from r-Java 2.0 calculations. There exists, however, three different types of underlying code to perform these calculations. Each code is suitable for a different scenario. Different codes and their inputs can be selected by clicking the "Code" tab in the right-most panel of the main window. For each code in r-Java 2.0, in order to run the code the user must click the "Calculate" button found in the toolbar or on the desktop.

- General Parameters: The parameters found on the "General" tab affect the r-process network calculations.
  - MF Cutoff: This is the lower limit in mass fraction that will 0 be considered non-zero during an r-process simulation. Lowering this cutoff increases accuracy, but at the expense of calculation speed.
  - Min. temp (WP): This is the lower limit threshold for 0 temperature during a Waiting Point network r-process calculation. Once the temperature drops below this threshold the r-process simulations stops.
  - Min.  $n_n$  (WP): This is the lower limit threshold for neutron Ο density during a Waiting Point network r-process calculation. Once the neutron density drops below this threshold the r-process simulations stops.

Code	
MF Cutoff:	1E-15 🖨
Min. temp (WP):	2E9 🖨
Min. n <sub>n</sub> (WP):	1E20 🖨
Min. Y /Y :	1 🖨
Min. Y <sub>n</sub> /Y;	1

- Min. Yn/Yr: This is the lowest neutron-to-r-process abundance ratio allowed for an r-0 process simulation. Once the ratio drops below this threshold the r-process simulations stops.

Note: Any changes that are made to the code (or general) parameters while a simulation is running will be immediately captured any incorporated into the current simulation. For example changing the environment type will change the way density is evolved. This can cause r-Java 2.0 to crash or yield unexpected results.

## NSE

NSE is short for "Nuclear Statistical Equilibrium" and represents a state where the rate of all forward nuclear reactions is equal to the rate of the reverse reactions. This code does not represent the conditions existing during the neutron matter decompression that r-Java is used to investigate, but is included as a "benchmark" or test of the underlying code. This being said, it is still useful for NSE work.

The code parameters tab in the "Code" panel displays the options for the NSE code. These options include:

Гуре:	NSE	V
<b>Т (К)</b> :		3E9
p (g cm <sup>-3</sup> ):		1E7
Ye:		0.3
		Coulomb interaction: 🗸
		Set default Abd.

Figure 8: Parameters available to the NSE code type

- **T**: The temperature of the system in Kelvin
- **q**: The density of the system in **g** cm<sup>-3</sup>
- Ye: The electron fraction
- Include coulomb interaction (yes/no): See "*r-Java 2.0 the nuclear physics*" for a discussion of this option.

After running the NSE code the user has the option to set the results from the NSE simulation as the initial mass fractions for an r-process simulation. To do this the user simply clicks the "Set initial abd."after having run the NSE code.

## **Fission**

The fission code calculates the fission fragmentation for a chosen nucleus. In order to add a target nucleus the user must click the "blue plus" symbol at the bottom of the fission parameters panel. Clicking the "blue plus" opens a window that lists all the available nuclei. Clicking on a nuclei and hitting "OK" adds that nucleus as a target for neutroninduced fission. The parameters that affect the mass fragmentation distribution are:

• **n Energy (MeV)**: The energy of the incident neutron in MeV.



- Standard I  $\delta V$ : The depth of the standard I fission fragmentation channel.
- **Standard II δV**: The depth of the standard II fission fragmentation channel.
- Standard I C: The strength of the standard I fission fragmentation channel.
- Standard II C: The strength of the standard II fission fragmentation channel.

For details on the standard channels and their associated parameters please see "r-Java 2.0 - the nuclear physics".

Once the parameters are chosen and the user clicks "Calculate" either the mass fragmentation will be plotted in the "Graphing" module or a message will appear in the "Messages" module explaining what the minimum incident neutron energy is required to induce fission.

## **r-Process**

The r-process code is the main code in r-Java 2.0. This code calculates the r-process abundance for the given initial conditions. The user is afforded many options when running r-process simulations with r-Java 2.0

#### **Network Type**

- Waiting point: This network calculation type assumes that the system is in  $(n, \gamma) \leftrightarrow (\gamma, n)$  equilibrium and thus the relative abundance along isotopic chains (nuclei with the same Z) is determined by neutron separation energy of each nucleus, neutron density and temperature. The reactions that are considered are;  $\alpha$ -decay,  $\alpha$ -capture,  $\beta$ -decay, neutron decay and fission. This network is faster to run than the full network, but is limited to high temperature (T > 2 10<sup>9</sup> K) and high neutron density (n<sub>n</sub> > 10<sup>20</sup> cm<sup>-3</sup>).
- Full network: This network calculation considers all the possible reactions (α-decay, α-capture, β-decay, β-delayed neutron emission, neutron decay, neutron capture, photo-dissociation and fission) for each nucleus in the system. This calculation runs slower than the waiting point approximation, but is not limited to the high temperature, high neutron density regime.

### **Setting Initial Abundances**

r-Java 2.0 is equipped with several different options for setting the initial abundance for an r-process simulation.

- Use NSE module: Once an NSE simulation is run the user can click the "Set initial abd." at the bottom of the NSE code panel. This will automatically set the resultant abundances from the NSE run as the initial abundances for the next r-process simulation.
- Import initial mass fractions: Within the "Data" module at

the bottom, directly adjacent to the table is where one can import the initial mass fractions that will be used in the next r-process simulation. The format of the mass fraction import file is; Z, A, mass fraction, each separated by the chosen character.

- Do not include the mass fraction of neutrons in the import file as they will be calculated by r-Java 2.0.
- In order for an r-process simulation to start the neutron-to-seed abundance ratio cannot be lower than the user defined minimum cut-off. Where the seed abundance is calculated as the sum of Y = (Mass Fraction)/A for all the initial mass fractions.
- Use the Data table: The user can as well simply enter the default mass fraction into the "Initial MF" column of the table seen on the "Data" Module. Note the neutron-to-seed ratio requirement specified above.
- Set initial element [WAITING POINT ONLY]: If the user is running an r-process simulation using the waiting point network there is a further option for setting the initial abundances. The user can specify Z<sub>0</sub> as the initial element and Y<sub>e</sub> as the initial electron fraction. r-Java 2.0 then calculates the initial abundances of the isotopes of Z<sub>0</sub>. In order to disable this option the user must set Z<sub>0</sub> to -1.
  - NOTE: for all other initial abundance methodologies the initial electron fraction is calculated by r-Java 2.0 using the initial abundances.

Import Initial Mass Fraction
Filename:
Column Separator
🔘 Tab
Space
O Other
Import

Fig. 10: Importing initial mass fractions.

🦲 m	assIm	portTest.t	xt - No	tepad
<u>F</u> ile	<u>E</u> dit	F <u>o</u> rmat	View	<u>H</u> elp
3 26 92		7 70 235	0. 0. 0.	2  1 1
		235	0.	-

Fig. 11: Example of initial mass fraction import file.

### **Environment Type**

The choice of astrophysical environment for the r-process determines how the physical parameters (namely temperature and density) of the system evolve. r-Java 2.0 gives the user a choice of three different possible astrophysical r-process sites and the option to define a custom density evolution profile. For details on the physics underlying each astrophysical environment please see "*r-Java 2.0 – the astrophysics*". Regardless of the environment type all r-process simulations have the following common parameters:

- **Duration [s]**: How long the r-process simulation is run for in seconds.
- **Decay time [s]**: How long the simulation will decay back to stability after the r-process stops (default is the age of the Universe) in seconds.
- **Update**: This specifies how many steps in the calculation before the graphs are updated. Setting this to a lower number allows you to see how the abundances are changing, but may take longer to calculate.

On top of these common parameters, each environment comes with a specific set of input parameters.

#### **High-Entropy Winds**

The parameters available in the high-entropy wind environment are:

- **T**<sub>0</sub> **[K]**: The initial temperature of the system in K
- **Entropy**: The entropy of the wind.
- V Exp [km/s]: The expansion speed of the wind in kilometers per second
- **R**<sub>0</sub> **[km]**: The initial radius of the wind packet in kilometers.

#### **Neutron Star Mergers**

The parameters available to the neutron star merger environment are:

eneral	
ode	
pe.	I-PTOCESS
nput Parameters	s 
Network:	Full Network
Environment	High-entropy winds
Т <sub>0</sub> (К):	3.0000000E9
p <sub>o</sub> (g cm <sup>-3</sup> ):	1.0000000E11
p (t):	Edit
Ye:	0.3000000
r [s]	0.10000000
Duration [s]:	1.0000000
Decay time [s]:	4.35400000E17
Z <sub>o</sub> :	-1 🖉
Update:	500 🖨
Z <sub>o</sub> : Update:	-1
ligh-entropy win 	ds 🖉
=ntropy:	250 🟹
V Exp (km/s):	7500 🖨
R <sub>o</sub> (km):	130 🖨

Fig. 12: Parameters available for the high-entropy wind environment.

- $T_0$  [K]: The initial temperature of the system in K
- ρ<sub>0</sub> [g cm<sup>-3</sup>]: The initial density of the system in g cm<sup>-3</sup>
- **Polytropic Index:** The neutron star matter is approximated by a polytrope and as such the equation of state is  $P = K\rho^{(n+1)/n}$  where **n** is the polytropic index.
- **P0:** The initial internal pressure of the system in units of MeVfm<sup>-3</sup>
- **V** Exp [km/s]: The expansion speed of the ejecta in kilometers per second
- **R**<sub>0</sub> **[km]**: The initial radius of the ejecta packet in kilometers.

#### **Quark Nova**

When the quark nova is selected as the r-process site, the underlying parameters (such as temperature and density) are then calculated based on the parameters listed below:

- $\mathbf{M}_{\star}$  [ $\mathbf{M}_{\odot}$ ]: The mass of the neutron star in solar masses
- **R**<sub>\*</sub> [km]: The radius of the neutron star in kilometers
- $\mathbf{M}_{ejecta}$  [ $\mathbf{M}_{\odot}$ ]: The mass of the neutron star ejecta
- Zeta: The percentage of quark nova energy transformed into kinetic energy of the ejecta.

Once the quark nova parameters are specified the user must click the red "Import" button in order to calculate the predicted physical parameters.

#### **Custom Environment**

Choosing the custom environment allows the user to define a custom density evolution profile. The parameters available are:

• **T**<sub>0</sub> **[K]**: The initial temperature in Kelvin.

General	
Code	
ype:	r-Process
Input Parameters	
Network:	Full Network
Environment:	Neutron star merger 🛛 🔻
T <sub>0</sub> [K]:	3.0000000E9
ρ <sub>0</sub> [g cm <sup>-3</sup> ]:	1.0000000E11
ρ (t):	Edit
Ye:	0.3000000
T [S]	0.1000000
Duration [s]:	1.0000000
Decay time [s]:	4.35400000E17
Z <sub>0</sub> :	-1
Update:	500
Neutron Star Mer	ger
P <sub>o</sub> [MeV fm <sup>-3</sup> ]	0.0005
V <sub>exp</sub> [km/s]:	1E5
R. [km]:	2

Fig. 13: Parameters available for the neutron star merger environment.

	-
eneral	
Code	
ype:	r-Process
Input Paramete	rs
Network:	Full Network
Environment:	Quark Nova 🛛 🔻
т <sub>о</sub> [K]:	3.0000000E9
ρ <sub>0</sub> [g cm <sup>-3</sup> ]:	1.0000000E11
ρ (t):	Edit
Ye:	0.3000000
T [S]	0.1000000
Duration [s]:	1.0000000
Decay time [s]:	4.35400000E17
Z <sub>o</sub> :	-1
Update:	500
Quark Nova —	
M <sub>★</sub> [M <sub>☉</sub> ]:	1.4
R <sub>*</sub> [km]:	10
M <sub>ejecta</sub> [M <sub>☉</sub> ]:	0.001
7.4	0.01

- 🗅 🗙

- ρ<sub>0</sub> [g cm<sup>-3</sup>]:: The initial density of the system in g cm<sup>-3</sup>
- $\rho(t)$ : The density evolution function, for static case set to  $\rho_0$ .
- $\tau$  [s]: Expansion timescale in seconds.

# Graphs

raphs are the main component by which the abundance information is conveyed to the user. Several types of graphs are available for plotting. The graphs are displayed in the "Graphs" module in the "Main Panel". By default one graph is set to fill the entire "Graph" module window, however one can choose to arrange the graphs in tile or cascade by clicking the "Tile" or "Cascade" button in the graph module toolbar (directly below the main r-Java 2.0 toolbar when the "Graph" module is attached to the main r-Java 2.0 window).

Individual graph properties can be adjusted by right clicking anywhere on the graph. Here you can save the graph information as a text file, or save it as an image. Additionally you can edit the "properties" of the graph. Many aesthetic options are available so that you can customize your graph any ways you like.

Each graph can be selected by clicking on the appropriate graph icon in the "Open Graphs" panel along the left-hand side of the "Graphs Module". Once a graph is selected, its parameters will become available in the "Graph Parameters" tab in the panel directly to the right adjacent of the graph.

## **Graph Parameters**

Within the "Graph Parameters" panel are the options for one to adjust the axes properties, modify the way the data sets are plotted and more. The options are divided into the following tabs:

- X axis / Y axis: These tabs allow the user to do such things as assign a variable to that axis, set the maximum and minimum as well as lock the axes (will not auto-adjust to fit data) and choose a log scale.
- **Data:** This tab provides the user with the options to change properties such as the line colour, width and style as well as scale the data.

## **Storing Plots**

When an r-process simulation is running the data will be plotted under the name "Primary Data [Point]" which appears in the window at the top of the "Data" tab in the "Graph Parameters" panel. The user can store graphed data by simply changing its name. This can be done during a simulation run without interrupting the code by changing the "Primary Data [Point]" name to anything else. A "Primary Data [Point]" data set will



Fig. 15: Open graphs panel.

reappear the next time the code reaches a time-step that is a multiple of the "Update" parameter set by the user (see Chapter 5, section "r-Process", sub-section "Environment Type").

## **Graph Types**

Several types of graphs are available including:

- Abundance: This is the main graph type in r-Java 2.0 and plots attributes of all the loaded nuclei. The variables plotted on the axes can be selected via the parameters in the "Graph Parameters" tab in "Parameters" the panel. plot Attributes you can include Z, A, N, Mass, Mass Fraction, abundance and even solar abundances.
- **Time**: This graph allows you to plot values for every time step in the r-process code calculation. The axis variables available to this graph type include; time, temperature, Q



Fig. 16: Graph module with the Data tab open in the Graph Parameters panel.

heat, matter density, entropy, electron fraction and neutron density.

• Rates: This graph will plot a temperature dependent rate of one or more isotopes versus temperature.